

# "First Science Run at Livingston LIGO Observatory: Data calibration and stability"

## Gabriela González Louisiana State University





# S1 run: Aug 23-Sept 9



## Lots of data...



#### http://www.phys.lsu.edu/faculty/gonzalez/S1LockStats/







# **Calibration? What calibration?**



# A "simple" DARN model



Rana Adhikari's Simulink model

# Why does calibration matter?



## **Control room calibration**



Why would calibration change?If 
$$C(f) \rightarrow \alpha C(f)$$
, $AS\_Q \rightarrow X_{ext} \frac{\alpha C(f)}{1 + \alpha H(f)}$ . $G(f) \rightarrow \beta G(f)$  $AS\_Q \rightarrow X_{ext} \frac{\alpha C(f)}{1 + \alpha \beta H(f)}$ .

We can get information about a from changes in the amplitude of known displacements: we push the mirrors with sine waves, or "calibration lines".



### How much did the calibration change?





### Not really :

gain is not the only indicator of noise (or even a good one)





#### There's a lot of work to do!

- Understand the noise:
  - Dominant sources
  - Not-so-dominant sources
  - Model comparison
  - ...
- Find out a good tracking calibration method
- Keep the alignment controlled so that the calibration does not change!
- Find good criteria to evaluate performance in REAL TIME