



How much does the calibration fluctuate (S2)?

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 α (t): depends on alignment, input power, offsets in error points...

 β (t): digital gain: fixed in S2, dynamic in S3 to keep $\alpha\beta$ "constant"

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LIGO 5% error in optical gain => freq. dep. error in calibration



4/23/200

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LIGO Measures of optical gain



- -- Amplitude of calibration line (SenseMon, LineMon, demodulation)
- $\stackrel{--}{\sim} \alpha \propto J_0 J_1 \propto \sqrt{(P_x + P_y)P_{sb}} \propto \sqrt{(QPDX + QPDY)SPOB}$





S2 L1 α, β





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S2 L1 α, β





S2 H1 α, β





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LIGO **S3 calibration : dynamic** β



A dynamic β : calculated at 16384Hz, low pass filtered at ~0.5 Hz, Read at ~ 8 Hz(?), written at 16 Hz, averaged by SenseMon at 1/60s



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"Fast" fluctuations: S2 L1 examples







Rough consistency between line amplitude and power function

Freq Resp: pole in power function? (not always there?)





Lowest cal line amplitude (10% error in minute time scale): no information on fast fluctuations from demod line, similar information from pwr function (less useism, same 1-2 Hz)











Large (not largest) line amplitude

-Consistent low freq amplitude

Large (too large?)1-2 Hz fluctuations







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Data-734247221-734250820 .1.1



Same as before, filtered below 0.5 Hz.

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Scaled by the reference value at the same time than for the SenseMon line, long time trends are consistent with demodulation, not so clear with pwr function.

Conclusions



- Fast fluctuations can be 10-15% pp, mostly at 0.1-0.2 Hz and 1-2 Hz
- 0.1-0.2 Hz fluctuations may be reconstructed from Power Function, not clear for long trends or for 1-2 Hz
- Fluctuations can be measured from demodulated line IF amplitude is large enough.
- Estimates on magnitude of fluctuations (but probably not measurements of fluctuations) may be obtained from blrms tools or seismometer channels.
- Lots of work!